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TI Branching fractions in the dissociative recombination of  $\text{NH}_4^+$  and  $\text{NH}_2^+$  molecular ions

SO ASTRONOMY AND ASTROPHYSICS

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DE ISM : molecules; ISM : clouds; molecular processes

ID HIGH-RESOLUTION MEASUREMENT; ABSOLUTE CROSS-SECTIONS;  
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AB Branching fractions in the dissociative recombination of  $\text{NH}_4^+$  and  $\text{NH}_2^+$  molecular ions with electrons were measured using the CRYRING heavy ion storage ring. We have determined complete branching fractions for  $\text{NH}_4^+$  at 0 eV and 2 meV collision energies, and at 0 eV collision energy for  $\text{NH}_2^+$ . We found the dissociative recombination of  $\text{NH}_4^+$  to be dominated by the two body, 'ammonia' channel. The branching fractions we obtained at 0 eV are 0.69 +/- 0.03 for the  $\text{NH}_3 + \text{H}$  channel, and for the breakup into the  $\text{NH}_2 + \text{H}_2$  and  $\text{NH}_2 + 2\text{H}$  channels 0.10 +/- 0.02 and 0.21 +/- 0.03, respectively. The values we obtained for 2 meV are about the same as at 0 eV. For the  $\text{NH}_2^+$  ion we obtained 0.66 +/- 0.01 for the three body  $\text{N} + \text{H} + \text{H}$  channel, 0.34 +/- 0.02 for  $\text{NH} + \text{H}$ , and no breakup into the  $\text{N} + \text{H}_2$  channel. Dissociative recombination of  $\text{NH}_4^+$  and  $\text{NH}_2^+$  is important as a potential source of some neutral molecules found in the interstellar clouds ( $\text{NH}_3$ ,  $\text{NH}_2$ , and  $\text{NH}$ ) and the measured branching fractions have important implications for modelling the chemistry of these clouds. We report also on the relative dissociative recombination cross sections for  $\text{NH}_4^+$  and  $\text{NH}_2^+$  for collision energies below 60 eV.

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