

# Excitation and fragmentation of $C_nN$ molecules; Fundamental aspects and application to astrochemistry.

T. Mahajan<sup>1\*</sup>, K. Béroff<sup>1</sup>, M. Chabot<sup>2</sup>, T. Pino<sup>1</sup>, T. K. C. Le<sup>1</sup>, M. Bonnin<sup>1</sup>, G. Martinet<sup>2</sup>, S. Bouneau<sup>2</sup>, L. Perrot<sup>2</sup>, A. Le Padellac<sup>3</sup>, F. Geslin<sup>2</sup>, N. De Sereville<sup>2</sup>, F. Hammache<sup>2</sup>, T. Launoy<sup>4</sup>.

<sup>1</sup>*Institut des Sciences Moléculaires d'Orsay (France).*

<sup>2</sup>*Institut de Physique Nucléaire d'Orsay (France).*

<sup>3</sup>*Centre d'Etude Spatiale des Rayonnements, Toulouse (France).*

<sup>4</sup>*Chimie Quantique et photophysique, Université libre de Bruxelles (Belgique).*

\**thejus.mahajan@u-psud.fr*

Amongst the ~180 molecules detected in the interstellar medium (ISM), carbon-based molecules are dominant, associating carbon atoms with hydrogen, nitrogen or oxygen ones.  $C_nN$  molecules, on the neutral but also on the anionic forms, have been detected in ISM [1] and have also been found in planetary atmospheres such as Titan, the largest satellite of Saturn [2].

In the astrochemistry field, dedicated to the study of the molecular composition and evolution in astrophysical environments, a strong demand emerges for molecular fragmentation data to be included in databases such as the recent KIDA (Kinetic Database for Astrochemistry) [3]. Using the dedicated AGAT setup nearby Tandem-Alto accelerator in Orsay, the AGAT collaboration was able to measure fragmentation branching ratios (BR's) for numerous carbon clusters and hydrocarbon molecules of various sizes and charges [4].

We recently began measurements with  $C_nN^+$  molecules. Apart from fragmentation BR's, collisional cross sections for electronic excitation, ionization and charge transfer in 2.25 a.u  $C_nN^+$  - He collisions have also been measured which are of fundamental interest. I will present first results obtained with the  $C_nN^+$  projectiles for collisional cross sections and associated fragmentation BRs and will compare to data obtained previously with pure carbon clusters.

## References

- [1] M. Agundez and V. Wakelam Chem. Rev **113** (2013) 8710.
- [2] V. Vuitton *et al.* Planetary and Space Science **57** (2009) 1558.
- [3] <http://kida.obs.u-bordeaux1.fr/>
- [4] M. Chabot *et al.* ApJ (2013) 771:90 and references therein.